Lecture 4: Ornstein-Zernike Formula

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In this lecture we see ...

- The mean-field theory discussed in the previous section does not tell us about the spatial correlation.
- In the previous lecture, we derived the continuous version of the Ising model, i.e., ϕ^4 model.
- We can apply the GBF variational approximation to the ϕ^4 Hamiltonian, with the variational Hamiltonian that has a non-trivial spatial structure.
- As a result, we obtain the Ornstein-Zernike form for the correlation function.

Variational approximation to ϕ^4 model

- Similar to the Ising model, generally it is impossible to obtain the exact solution of ϕ^4 model by analytical means. So, we need some approximation. The simplest mean-field approximation neglecting the special variance results in the same type of
- We will first move to the momentum space.
- Then, we will apply the GBF variational principle by taking the Gaussian theory as the trial Hamiltonian.
- As a result, we will obtain the mean-field evaluation of the spatial correlation function, which is called Ornstein-Zernike form.

Switching to the momentum space

Starting from ϕ^4 model in the discrete space,

$$
\mathcal{H}=a^d\sum_{\bm{x}}\left(\rho|\nabla\phi_{\bm{x}}|^2+t\phi_{\bm{x}}^2+u\phi_{\bm{x}}^4-h\phi_{\bm{x}}\right),\,
$$

by Fourier transformation $\phi_{\bm{x}} = L^{-d} \sum_{\bm{x}} \bm{x}$ k $e^{i\bm{k}\bm{x}}\tilde{\phi}_{\bm{k}},$ we obtain

$$
\mathcal{H} = \frac{1}{L^d} \sum_{\mathbf{k}} (\rho k^2 + t) |\tilde{\phi}_{\mathbf{k}}|^2 \n+ \frac{u}{L^{3d}} \sum_{\mathbf{k}_1 \cdots \mathbf{k}_4} \delta_{\sum_{\mu=1}^4 \mathbf{k}_\mu, \mathbf{0}} \tilde{\phi}_{\mathbf{k}_1} \tilde{\phi}_{\mathbf{k}_2} \tilde{\phi}_{\mathbf{k}_3} \tilde{\phi}_{\mathbf{k}_4} - h \tilde{\phi}_{\mathbf{0}}.
$$
\n(1)

(If you prefer continuous wave numbers, you could instead use $\mathcal{H} = \int \frac{d^d \mathbf{k}}{(2\pi)^d}$ $\frac{d^d\bm{k}}{(2\pi)^d}$ $(\rho k^2 + t)\tilde{\phi}^*_{\bm{k}}\tilde{\phi}_{\bm{k}} + u\int \frac{d^d\bm{k}_1\cdots d^d\bm{k}_4}{(2\pi)^{4d}}$ $\frac{\bm{k}_1\cdots d^d\bm{k}_4}{(2\pi)^{4d}}\, \delta\left(\sum_\mu \bm{k}_\mu\right)\, \tilde{\phi}_{\bm{k}_1}\cdots \tilde{\phi}_{\bm{k}_4} - h \tilde{\phi}_{\bm{0}}\, \, .)$

Supplement: Convention (Fourier transformation)

In this lecture, we use the following conventions:

$$
a = (\text{lattice constant}), \quad L = (\text{system size}), \quad N \equiv \frac{L^d}{a^d} = (\# \text{ of sites})
$$
\n
$$
\tilde{\phi}_{\mathbf{k}} = \int_0^L d^d \mathbf{x} \, e^{-i\mathbf{k}\mathbf{x}} \phi_{\mathbf{x}} = a^d \sum_{\mathbf{x}} e^{-i\mathbf{k}\mathbf{x}} \phi_{\mathbf{x}}
$$
\n
$$
\phi_{\mathbf{x}} = \int_{-\pi/a}^{\pi/a} \frac{d^d \mathbf{k}}{(2\pi)^d} e^{i\mathbf{k}\mathbf{x}} \tilde{\phi}_{\mathbf{k}} = L^{-d} \sum_{\mathbf{k}} e^{i\mathbf{k}\mathbf{x}} \tilde{\phi}_{\mathbf{k}}
$$

The tilde $\tilde{ }$ is often dropped when there is no fear of confusion.

$$
G(\mathbf{x}', \mathbf{x}) \equiv \langle \phi_{\mathbf{x}'} \phi_{\mathbf{x}} \rangle, \quad G_{\mathbf{k}', \mathbf{k}} \equiv L^{-d} \langle \phi_{\mathbf{k}'} \phi_{\mathbf{k}} \rangle
$$

For translationally and rotationally symmetric case,

 $G(\boldsymbol{x}',\boldsymbol{x}) = G(|\boldsymbol{x}'-\boldsymbol{x}|),\quad G_{\boldsymbol{k}',\boldsymbol{k}} = \delta_{\boldsymbol{k}'+\boldsymbol{k},\boldsymbol{0}}G_{|\boldsymbol{k}|},\quad G_{|\boldsymbol{k}|} \equiv L^{-d}\langle|\phi_{\boldsymbol{k}}|^2\rangle$

GBF variational approximation (1)

Let us consider a trial Hamiltonian with variational parameter $\epsilon_{\mathbf{k}}$,

$$
\mathcal{H}_0 \equiv \frac{1}{L^d} \sum_{\mathbf{k}} \epsilon_{\mathbf{k}} |\phi_{\mathbf{k}}|^2
$$
\n
$$
Z_0 = \int D\phi \, e^{-\frac{1}{L^d} \sum_{\mathbf{k}} \epsilon_{\mathbf{k}} |\phi_{\mathbf{k}}|^2} = \prod_{\mathbf{k}} \zeta_{\mathbf{k}}
$$
\n
$$
\langle |\phi_{\mathbf{k}}|^2 \rangle_0 = \frac{L^d}{2\epsilon_{\mathbf{k}}}, \quad \zeta_{\mathbf{k}} \equiv \left(\frac{\pi L^d}{\epsilon_{\mathbf{k}}}\right)^{1/2}
$$
\n
$$
E_0 = \frac{1}{L^d} \sum_{\mathbf{k}} \epsilon_{\mathbf{k}} \langle |\phi_{\mathbf{k}}|^2 \rangle_0 = \sum_{\mathbf{k}} \frac{\epsilon_{\mathbf{k}}}{L^d} \frac{L^d}{2\epsilon_{\mathbf{k}}} = \sum_{\mathbf{k}} \frac{1}{2} = \frac{N}{2} \quad \text{``Equipartition''}
$$
\n
$$
-TS_0 = F_0 - E_0 = -\sum_{\mathbf{k}} \frac{1}{2} \log \frac{\pi L^d}{\epsilon_{\mathbf{k}}} = \frac{1}{2} \sum_{\mathbf{k}} \log \epsilon_{\mathbf{k}}
$$

(Additive constants have been omitted.)

GBF variational approximation (2)

$$
\langle \mathcal{H} \rangle_0 = \frac{1}{L^d} \sum_{\mathbf{k}} (\rho k^2 + t) \langle |\phi_{\mathbf{k}}|^2 \rangle_0 + \frac{u}{L^{3d}} \sum_{\mathbf{k}_1 \cdots \mathbf{k}_4} \delta_{\sum \mathbf{k}, \mathbf{0}} \langle \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \phi_{\mathbf{k}_3} \phi_{\mathbf{k}_4} \rangle_0
$$

=
$$
\frac{1}{L^d} \sum_{\mathbf{k}} (\rho k^2 + t) \langle |\phi_{\mathbf{k}}|^2 \rangle_0 + \frac{3u}{L^{3d}} \sum_{\mathbf{k}, \mathbf{k'}} \langle |\phi_{\mathbf{k}}|^2 \rangle_0 \langle |\phi_{\mathbf{k'}}|^2 \rangle_0 \quad \text{(Wick)}
$$

We have used $\langle \phi_{k'} \phi_k \rangle_0 = \delta_{k',-k} \langle |\phi_k|^2 \rangle_0.$ In terms of $G_{\bm{k}}\equiv L^{-d}\langle|\phi_{\bm{k}}|^2\rangle_0=(2\epsilon_{\bm{k}})^{-1}$, we obtain

$$
F_{\mathsf{v}} = \langle \mathcal{H} \rangle_0 - TS_0 = \sum_{\mathbf{k}} (\rho k^2 + t) G_{\mathbf{k}} + \frac{3u}{L^d} \left(\sum_{\mathbf{k}} G_{\mathbf{k}} \right)^2 + \frac{1}{2} \sum_{\mathbf{k}} \log \epsilon_{\mathbf{k}}
$$

Thus we have, $f_{\mathsf{v}} \equiv L^{-d} F_{\mathsf{v}} = B + 3uA^2 + \frac{1}{2L}$ $2L^d$ \sum k $\log \epsilon_{\boldsymbol{k}},$ (3)

$$
\text{where}\quad A\equiv \frac{1}{L^d}\sum_{\boldsymbol{k}}G_{\boldsymbol{k}},\text{ and }B\equiv \frac{1}{L^d}\sum_{\boldsymbol{k}}(\rho k^2+t)G_{\boldsymbol{k}}.
$$

Stationary condition

$$
0 = \frac{\partial F_v}{\partial \epsilon_k} = (\rho k^2 + t + \sigma) \frac{\partial G_k}{\partial \epsilon_k} + \frac{1}{2\epsilon_k}
$$

\n
$$
\left(\sigma \equiv 6uA = \frac{6u}{L^d} \sum_k G_k\right) \qquad \cdots \text{ Spatial fluctuation shifts}
$$

\n
$$
= (\rho k^2 + t + \sigma) \left(-\frac{1}{2\epsilon_k^2}\right) + \frac{1}{2\epsilon_k}
$$

\n
$$
\Rightarrow \epsilon_k = \rho k^2 + t + \sigma = \rho(k^2 + \kappa^2) \qquad \left(\kappa \equiv \sqrt{\frac{t + \sigma}{\rho}}\right)
$$

Ornstein-Zernike form $G_k \propto$ 1 $\frac{1}{k^2 + \kappa^2}, \quad \kappa \propto \sqrt{T-T_c}$

Supplement: Wick's theorem

Theorem 1 (Wick)

When the distribution function is Gaussian, any multi-point correlator factorizes in pairs.

Example 2 (4-point correlator)

Ex: When the Hamiltonian is $\mathcal{H}=\frac{1}{2}$ $\frac{1}{2} \boldsymbol{\phi}^\mathsf{T} A \boldsymbol{\phi}$ with A being a real positive-definite symmetric matrix,

 $\langle \phi_1 \phi_2 \phi_3 \phi_4 \rangle = \langle \phi_1 \phi_2 \rangle \langle \phi_3 \phi_4 \rangle + \langle \phi_1 \phi_3 \rangle \langle \phi_2 \phi_4 \rangle + \langle \phi_1 \phi_4 \rangle \langle \phi_2 \phi_3 \rangle$ $=\Gamma_{12}\Gamma_{34}+\Gamma_{13}\Gamma_{24}+\Gamma_{14}\Gamma_{23}$

where $\Gamma\equiv A^{-1}$ and $\langle\cdots\rangle\equiv$ $\int D\phi \, e^{-\mathcal{H}(\phi)} \cdots$ $\int D\phi \, e^{-\mathcal{H}(\phi)}$

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Supplement: Proof of Wick's theorem

If we define $\Xi \equiv \int D \bm{\phi} \, e^{-\frac{1}{2} \bm{\phi}^\mathsf{T} A \bm{\phi} + \bm{\xi}^\mathsf{T} \bm{\phi}},$ the correlation function can be expressed as its derivatives,

$$
\langle \phi_{k_1} \phi_{k_2} \cdots \phi_{k_{2p}} \rangle = \Xi^{-1} \left(\frac{\partial}{\partial \xi_{k_1}} \cdots \frac{\partial}{\partial \xi_{k_{2p}}} \Xi \right) \Big|_{\xi \to 0}.
$$

Now notice that $\Xi \propto e^{\frac{1}{2}}$ $\frac{1}{2} \boldsymbol{\xi}^\mathsf{T} \Gamma \boldsymbol{\xi}$, with $\Gamma \equiv A^{-1}$, which yields

$$
\Xi = 1 + \sum_{ij} \frac{\Gamma_{ij}}{2} \xi_i \xi_j + \frac{1}{2} \sum_{ij} \sum_{kl} \frac{\Gamma_{ij}}{2} \frac{\Gamma_{kl}}{2} \xi_i \xi_j \xi_k \xi_l + \cdots
$$

Therefore, the $2p$ -body correlation becomes

$$
\frac{1}{p!} \sum_{i_1 j_1} \sum_{i_2 j_2} \cdots \sum_{i_p j_p} \frac{\Gamma_{i_1 j_1}}{2} \frac{\Gamma_{i_2 j_2}}{2} \cdots \frac{\Gamma_{i_p j_p}}{2} \delta_{\{k_1, k_2, \cdots, k_{2p}\}, \{i_1, j_1, i_2, j_2, \cdots, i_p, j_p\}}
$$
\n
$$
= \sum \Gamma_{i_1 j_1} \Gamma_{i_2 j_2} \cdots \Gamma_{i_p j_p} \quad \text{(Summation over all pairings of } \{k_1, \cdots, k_{2p}\} \text{)}
$$

Real-space correlation function

$$
G_{\mathbf{k}} = L^{-d} \langle |\phi_{\mathbf{k}}|^2 \rangle = \frac{1}{2\epsilon_{\mathbf{k}}} = \frac{1}{2(\rho k^2 + t + \sigma)}
$$

\n
$$
G(\mathbf{x}' - \mathbf{x}) \equiv \langle \phi_{\mathbf{x}'} \phi_{\mathbf{x}} \rangle = L^{-2d} \sum_{\mathbf{k}, \mathbf{k}'} e^{i\mathbf{k}' \mathbf{x}'} e^{i\mathbf{k} \mathbf{x}} \langle \phi_{\mathbf{k}'} \phi_{\mathbf{k}} \rangle
$$

\n
$$
= L^{-d} \sum_{\mathbf{k}, \mathbf{k}'} e^{i\mathbf{k}' \mathbf{x}'} e^{i\mathbf{k} \mathbf{x}} \delta_{\mathbf{k}' + \mathbf{k}, 0} G_{\mathbf{k}} = L^{-d} \sum_{\mathbf{k}} e^{-i\mathbf{k}(\mathbf{x}' - \mathbf{x})} \frac{1}{2\epsilon_{\mathbf{k}}}
$$

\n
$$
G(\mathbf{x}) = \int \frac{d^d \mathbf{k}}{(2\pi)^d} \frac{e^{i\mathbf{k} \mathbf{x}}}{2\epsilon_{\mathbf{k}}} = \frac{1}{2} \int \frac{d^d \mathbf{k}}{(2\pi)^d} \frac{e^{i\mathbf{k} \mathbf{x}}}{\rho k^2 + t + \sigma}
$$

\n
$$
\int_{\infty}^{\infty} \frac{1}{\rho} \frac{\kappa^{d-2}}{(\kappa r)^{\frac{d-1}{2}}} e^{-\kappa r} \quad (\kappa r \gg 1, \ \kappa \equiv \sqrt{\frac{t + \sigma}{\rho}}) \quad (T > T_c)
$$

\n
$$
T = T_c)
$$

(∗ · · · see supplement)

Mean-field values of ν and η

$$
G(r) \sim \begin{cases} \frac{1}{\rho} \frac{\kappa^{d-2}}{(\kappa r)^{\frac{d-1}{2}}} e^{-\kappa r} & (\kappa r \gg 1, \ \kappa \equiv \sqrt{\frac{t+\sigma}{\rho}}) & (T > T_c) \\ \frac{1}{\rho} \frac{1}{r^{d-2}} & (T = T_c) \end{cases}
$$

Mean-field value of ν

$$
\text{For } T > T_c, \quad G(r) \propto \frac{1}{r^{\frac{d-1}{2}}} e^{-r/\xi}, \quad \xi = \kappa^{-1} \propto \frac{1}{\sqrt{|T - T_c|}} \Rightarrow \nu_{\text{MF}} = \frac{1}{2}
$$

Mean-field value of η

At
$$
T = T_c
$$
, $G(r) \propto \frac{1}{r^{d-2}} \Rightarrow \eta_{MF} = 0$ $\left(G(r) \propto \frac{1}{r^{d-2+\eta}} \right)$

$$
\textsf{CF: } (\nu, \eta) = \left\{ \begin{array}{lll} (\quad \ \ 0.5 \quad \ \ , \quad 0 \quad \quad \)\,\, (d \geq 4) \\ (\quad \ \ 0.63002(10) \quad \ \ , \ \ 0.03627(10) \quad \)\,\, (d=3) \quad (PRB82(2010), 174433) \\ (\quad \ \ 1 \quad \quad \ \ , \ \ 0.25 \quad \quad \)\,\, (d=2) \end{array} \right.
$$

Supplement: Evaluation of the asymptotic form $(T>T_c)$

$$
\int d\mathbf{k} \frac{e^{i\mathbf{kx}}}{k^2 + \kappa^2} = \int d\mathbf{k} e^{i\mathbf{kx}} \int_0^\infty dt e^{-t(k + \kappa^2)}
$$

=
$$
\int_0^\infty dt e^{-t\kappa^2} \int d\mathbf{k} e^{-t k^2 + i\mathbf{k}}
$$

=
$$
\int_0^\infty dt e^{-t\kappa^2} \int d\mathbf{k} e^{-t(k - \frac{i}{2t}\mathbf{x})^2 - \frac{\mathbf{x}^2}{4t}} = \int_0^\infty dt \left(\frac{\pi}{t}\right)^{\frac{d}{2}} e^{-\kappa^2 t - \frac{r^2}{4t}}
$$

(Here we define x so that $t\equiv \frac{r}{2\kappa}x$ and $\kappa^2 t + \frac{r^2}{4t}$ $\frac{r^2}{4t} = \frac{\kappa r}{2}(x + x^{-1}).$

$$
= \int_0^\infty dx \left(\frac{\pi}{x}\right)^{\frac{d}{2}} \left(\frac{2\kappa}{r}\right)^{\frac{d}{2}-1} e^{-\frac{\kappa r}{2}(x+x^{-1})}
$$

(For $\kappa r \gg 1$, we use $x + x^{-1} \approx 2 + \epsilon^2$ where $\epsilon \equiv x - 1$.)

$$
\approx \pi^{\frac{d}{2}} \left(\frac{2\kappa}{r}\right)^{\frac{d}{2}-1} e^{-\kappa r} \left(\frac{2\pi}{\kappa r}\right)^{\frac{1}{2}} \sim \frac{\kappa^{d-2}}{(\kappa r)^{\frac{d-1}{2}}} e^{-\kappa r}
$$

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Supplement: Evaluation of the asymptotic form $(T=T_c)$

As before, we have

$$
\int d\mathbf{k} \, \frac{e^{i\mathbf{kx}}}{k^2 + \kappa^2} = \int_0^\infty dt \, \left(\frac{\pi}{t}\right)^{\frac{d}{2}} e^{-\kappa^2 t - \frac{r^2}{4t}}
$$

Here, by setting $\kappa=0$ $(T=T_c)$,

$$
= \int_0^\infty dt \left(\frac{\pi}{t}\right)^{\frac{d}{2}} e^{-\frac{r^2}{4t}}
$$
\n(By defining $\eta \equiv \frac{r^2}{4t}$)

\n
$$
= \left(\frac{r^2}{4}\right)^{1-\frac{d}{2}} \pi^{\frac{d}{2}} \Gamma\left(\frac{d}{2} - 1\right) \sim \frac{1}{r^{d-2}}
$$

Gaussian MF approximation below T_c (1)

 \bullet To deal with the spontaneous magnetization below T_c , we must introduce a symmetry-breaking field η as a new variational parameter,

$$
\mathcal{H}_0 = L^{-d} \sum_{\mathbf{k}} \epsilon_{\mathbf{k}} |\phi_{\mathbf{k}}|^2 - \eta \phi_{\mathbf{k} = \mathbf{0}}
$$

 \bullet It is, then, a little tedious but not hard to see that (3) is replaced by

$$
f_{\mathsf{v}} \stackrel{*}{=} B + tm^2 + u(3A^2 + 6Am^2 + m^4) + \frac{1}{2L^d} \sum_{\mathbf{k}} \log \epsilon_{\mathbf{k}}, \qquad \text{(4)}
$$

where $m \equiv \langle \phi_x \rangle_0$ and, as before,

$$
A\equiv\frac{1}{L^d}\sum_{\pmb{k}}\frac{1}{2\epsilon_{\pmb{k}}},\quad B\equiv\frac{1}{L^d}\sum_{\pmb{k}}\frac{\rho\pmb{k}^2+\pmb{t}}{2\epsilon_{\pmb{k}}}
$$

Gaussian MF approximation below T_c (2)

• From $\partial f_{\mathsf{v}}/\partial m = 0$, we obtain

$$
t + 6uA + 2um2 = 0
$$

or
$$
m2 = -\frac{t + \sigma}{2u} \quad (\sigma \equiv 6uA)
$$
 (5)

• From $\partial f_{\mathsf{v}}/\partial \epsilon_{\mathbf{k}} = 0$ $(\mathbf{k} \neq \mathbf{0})$, we obtain

$$
\epsilon_{\mathbf{k}} = \rho k^2 + t + 6u(A + m^2).
$$

Using (5),
$$
\epsilon_{\mathbf{k}} = \rho k^2 - 2(t + \sigma) = \rho (k^2 + \hat{\kappa}^2) \quad \left(\hat{\kappa}^2 \equiv \frac{-2(t + \sigma)}{\rho}\right)
$$

Thus, we have obtained the Ornstein-Zernike type Green's function

$$
G_k = \frac{1}{2\epsilon_k} = \frac{1}{2\rho(k^2 + \hat{\kappa}^2)} \quad (T < T_c)
$$

The correlation length is $1/\sqrt{2}$ √ 2 times smaller than the high- T side.

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$$
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$$

Supplement: Wick's theorem with symmetry-breaking field

For deriving (4), since the external field distorts the Gaussian distribution, which is the precondition to the Wick's theorem, we must apply the theorem to the fluctuation $\delta\phi_x \equiv \phi_x - \langle \phi_x \rangle_0$, not ϕ itself. In the momentum space, by defining $\delta\phi_{\bm k}\equiv\phi_{\bm k}-\bar\phi_{\bm 0}\delta_{\bm k}$ $(\delta_{\bm k}\equiv\delta_{\bm k,\bm 0},\,\bar\phi_{\bm 0}=L^dm)$,

$$
\langle \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \phi_{\mathbf{k}_3} \phi_{\mathbf{k}_4} \rangle_0
$$

=\langle (\bar{\phi}_0 \delta_{\mathbf{k}_1} + \delta \phi_{\mathbf{k}_1}) (\bar{\phi}_0 \delta_{\mathbf{k}_2} + \delta \phi_{\mathbf{k}_2}) (\bar{\phi}_0 \delta_{\mathbf{k}_3} + \delta \phi_{\mathbf{k}_3}) (\bar{\phi}_0 \delta_{\mathbf{k}_4} + \delta \phi_{\mathbf{k}_4}) \rangle_0
= \bar{\phi}_0^4 \delta_{\mathbf{k}_1} \delta_{\mathbf{k}_2} \delta_{\mathbf{k}_3} \delta_{\mathbf{k}_4} + \bar{\phi}_0^2 (\delta_{\mathbf{k}_1} \delta_{\mathbf{k}_2} \langle \delta \phi_{\mathbf{k}_3} \delta \phi_{\mathbf{k}_4} \rangle_0 + 5 \text{ similar terms})
+ (\langle \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \rangle_0 \langle \phi_{\mathbf{k}_3} \phi_{\mathbf{k}_4} \rangle_0 + 2 \text{ similar terms})

Therefore, we obtain

$$
\sum_{\mathbf{k}_1,\mathbf{k}_2,\mathbf{k}_3,\mathbf{k}_4} \delta_{\sum \mathbf{k}} \langle \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \phi_{\mathbf{k}_3} \phi_{\mathbf{k}_4} \rangle_0
$$
\n
$$
= \bar{\phi}_0^4 + 6 \bar{\phi}_0^2 \sum_{\mathbf{k}_1} \langle \delta \phi_{\mathbf{k}_1} \delta \phi_{-\mathbf{k}_1} \rangle_0 + 3 \sum_{\mathbf{k}_1,\mathbf{k}_3} \langle \phi_{\mathbf{k}_1} \phi_{-\mathbf{k}_1} \rangle_0 \langle \phi_{\mathbf{k}_3} \phi_{-\mathbf{k}_3} \rangle_0
$$

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Exercise 4.1: In the lecture, in obtaining the OZ form for the correlation function, we employed the variational Hamiltonian that has the form $L^{-d}\sum_{\bm{k}}\epsilon_{\bm{k}}|\phi_{\bm{k}}|^2-\eta\phi_{\bm{k}=0}.$ What if we used $\mathcal{H}_0=\lambda\sum_{\bm{x}}(\phi_{\bm{x}}-m)^2$ instead? (Here, λ and m are variational parameters.) Obtain the equations of state that relates λ and m to ρ , t and u. For the sake of simplicity, consider the case where $h = 0$.

[W](#page-7-0)e apply the GBF inequality $F_v \equiv F_0 + \langle H - H_0 \rangle_0 \geq F$ to

$$
\mathcal{H} = \sum_{\mathbf{x}} (\rho (\nabla \phi)^2 + t \phi^2 + u \phi^4 - h \phi)
$$

and

$$
\mathcal{H}_0 = \lambda \sum_{\boldsymbol{x}} (\phi - m)^2.
$$

For F_0 , we have

$$
\beta F_0/N = -\log \int d\phi e^{-\lambda(\phi - m)^2} = \frac{1}{2} \log(\lambda/\pi).
$$

For calculating $\langle\mathcal{H}\rangle_0$ and $\langle\mathcal{H}_0\rangle_0$, we apply Wick's theorem to $\langle(\delta\phi)^n\rangle_0$ with $\delta\phi\equiv\phi-m$ and $\langle(\delta\phi)^2\rangle_0=1/(2\lambda)$, to obtain,

$$
\langle \phi \rangle_0 = m, \quad \langle \phi^2 \rangle_0 = \langle (m + \delta \phi)^2 \rangle_0 = m^2 + \langle \delta \phi^2 \rangle_0 = m^2 + \frac{1}{2\lambda},
$$

$$
\langle \phi^4 \rangle_0 = \langle (m + \delta \phi)^4 \rangle_0 = m^4 + 6m^2 \langle \delta \phi^2 \rangle_0 + \langle \delta \phi^4 \rangle_0
$$

$$
= m^4 + 6m^2 \langle \delta \phi^2 \rangle_0 + 3 \langle \delta \phi^2 \rangle_0^2 = m^4 + \frac{3m^2}{\lambda} + \frac{3}{4\lambda^2},
$$

$$
\langle (\nabla \phi)^2 \rangle_0 = \frac{1}{2} \sum_{\delta} \langle (\phi_{r+\delta} - \phi_r)^2 \rangle_0 = \frac{z}{2} \langle \phi_{r+\delta}^2 + \phi_r^2 - 2\phi_{r+\delta}\phi_r \rangle_0 = \frac{z}{2\lambda} = \frac{d}{\lambda}.
$$

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Then, noting that β is included in the definition of the Hamiltonians and the free energy,

$$
\frac{F_v}{N} = \frac{1}{2}\log\frac{\lambda}{\pi} + \left(\frac{d\rho}{\lambda} + \frac{t}{2\lambda} + \frac{3u}{4\lambda^2}\right) + \left(t + \frac{3u}{\lambda}\right)m^2 + um^4 - hm - \frac{1}{2}
$$

From the stationary conditions, we obtain the mean-field type behaviors. Specifically, at $h = 0$,

$$
m = 0, \ \lambda = d\rho + t - t_c \quad (t > t_c)
$$

$$
m = \sqrt{(t_c - t)/6u}, \ \lambda = d\rho \quad (t < t_c)
$$

where $t_c \equiv -3u/\lambda$.

The same results can be obtained by working with the wave-number space instead of the real space. We substitute ϕ_k in ${\cal H}$ by $\phi_k\equiv\delta\phi_k+mL^d\delta_{k,0}$, and consider $\mathcal{H}_0 \equiv \lambda L^{-d} \sum_k \delta \phi_k.$ One thing that needs a careful treatment the discrete nature of " ∇ ". If we simply replace $(\nabla\phi)^2$ by $-k^2|\phi_k|^2$, as we usually do in the

wave-number space calculation, the result would be slightl]y diferent from the real-space calculation (though the diference is merely quantitative, and not so essential). Since ∇ here is the difference rather than the differentiation, to be precise, we must use $\sum_{\alpha=1}^d 2(1-\cos k_\alpha)$ in the place of $k^2.$ When we take the average of this term over the wave numbers in the 1st Brillouin zone, the cosine term yields zero, while the constant term yields $2d = z$, which is exactly the same as the coefficient ρ/λ term in the real-space calculation.

